QED second order corrections on the speed of light at low temperature

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Abstract

We want to study thermal corrections on the speed of light at low temperature considering temperature dependence of photon vacuum polarization tensor at two-loop level in the standard QED. It is found that the heat bath behaves as a dispersive medium for the propagation of light and reduces its speed proportional to the second order of temperature. Similarities and differences, in light of already known calculations which are based on Euler-Heisenberg Lagrangian and/or those using temperature dependent electromagnetic properties of the medium are discussed.

Keywords: QED radiative corrections; finite temperature quantum field theory; the speed of light

1. Introduction

The speed of light, as a fundamental constant of physics and as a speed independent of the motion of its source has a deep connection with the physical vacuum properties. The vacuum in quantum field theory is an ocean of particles that in an invisible time is created and then destroyed so that they are not detectable (virtual photons); it is a stormy sea of quantum fluctuations as the physical basis for the calculations of renormalization effects and radiation corrections. There are a number of theoretically well-known and experimentally confirmed phenomena as the spontaneous emission, and the Casimir effect which all originate from the quantum vacuum [1]. Renormalization process and radiative corrections are also due to quantum vacuum where, in QED, affect the physical (renormalized) mass of electron and its charge [2,3]. Considering the finite temperature quantum field theory, the propagators are affected by the presence of the background heat bath. The corresponding contributions are calculated either in Euclidean or Minkowski space using imaginary or real time formalism respectively [4]. In this paper, we use the real time formulation wherein energy is a continuous variable in conventional field theory which explicitly separates the $T = 0$ and $T \neq 0$ components, and is considerably simple [5]. The real time propagators at low temperatures (relative to the electron mass; $T << m_e$) can be written as [6]:

$$D_\beta^{\rho\sigma}(l) = g^{\rho\sigma}[\frac{1}{l^2 + i\varepsilon} - 2\pi i\delta(l^2)n_\beta(l)], \tag{1}$$

for photon propagator, where

$$n_\beta(l) = \frac{1}{e^{\beta|\vec{l}|} - 1} \tag{2}$$

is the Bose-Einstein distribution and

$$S_\beta(k) = \frac{(\not{k} + m)}{k^2 - m^2 + i\varepsilon}, \tag{3}$$

for fermions (here, the Fermi-Dirac statistical factor $n_F(E) = 1/(e^{\beta E} + 1)$ in the fermion propagator at finite temperature is ignorable [7]. Thus, at $T << m_e$, the hot fermions' contribution in background is suppressed and the only contribution is made by the hot photons.

In recent decades, several papers have calculated the thermal effects on the speed of light using effective (Euler-Heisenberg) Lagrangian [8,10] and/or standard QED radiative corrections on the electromagnetic properties of the medium. Unfortunately, both categories deal with some difficulties and we shall point out in section 4. Here, working technically in parallel and similar to [9], considering the possibility of photon-photon scattering and examining the temperature dependence of the vacuum polarization tensor based on standard QED radiative corrections we obtain the corresponding thermal correction, on the speed of light. It will be seen that "the finite temperature" (the heat bath) behaves as a
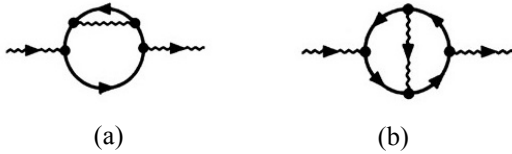
Figure 1. Two loop vacuum polarization.

dispersive medium for the propagation of light and reduces its speed proportional to the square form of temperature.

2. Radiative correction on the vacuum polarization tensor and the photon mass

The vacuum polarization tensor $\Pi_{\mu\nu}(p)$, as a second rank Lorentz tensor, can be constructed in terms of $g_{\mu\nu}, p_\mu p_\nu$, and scalar function p^2 as in the following [10]:

$$\Pi_{\mu\nu}(p) = D g_{\mu\nu} + g_{\mu\nu} p^2 \Pi^{(1)}(p^2) \qquad (4)$$
$$+ p_\mu p_\nu \Pi^{(2)}(p^2)$$

In the limit $p^2 \to 0$ and up to terms of higher order in α, this yields

$$\frac{-ig_{\alpha\beta}}{p^2 + i\varepsilon} \to \frac{-ig_{\alpha\beta}}{p^2 + i\varepsilon} + \frac{-ig_{\alpha\mu}}{p^2 + i\varepsilon} iDg^{\mu\nu} \frac{-ig_{\nu\beta}}{p^2 + i\varepsilon}$$
$$+ \frac{-ig_{\alpha\mu}}{p^2 + i\varepsilon} iDg^{\mu\nu} \frac{-ig_{\nu\sigma}}{p^2 + i\varepsilon} iDg^{\sigma\tau} \frac{-ig_{\tau\beta}}{p^2 + i\varepsilon} + \dots \quad (5)$$
$$\cong \frac{-ig_{\mu\nu}}{p^2 - D + i\varepsilon} \ .$$

This is just the propagator of a boson with mass $\sqrt{D}$ (a "heavy photon"). Thus, there is the following relation between the effective (dynamical) photon mass and the vacuum polarization tensor:

$$g^{\mu\nu}\Pi_{\mu\nu}(p,T) = 4D. \qquad (6)$$

2. 1. Two loop vacuum polarization

At one loop level, because of the absence of self-interaction of photons at the approximation of low temperature limit we are considering here, the vacuum polarization tensor (in α order) is zero.

At higher-loop level, the loop integrals involve a combination of temperature independent (cold) and temperature dependent (hot) terms which appear due to the overlapping propagator terms [11]. In order α^2, this contribution basically comes from self mass (in figure 1.(a) with counter term ⊸◯⊸) and vertex type (in figure 1. (b) with counter term ⊸◯⊸ or ⊸◯⊸) of electron loop corrections inside the vacuum polarization tensor [3]. The expression for two loop photon self energy in figure 1 (a) is

$$\Pi_{\mu\nu}^{(1)}(p) =$$
$$e^4 \int \frac{d^4k}{(2\pi)^4} \int \frac{d^4l}{(2\pi)^4} \text{Tr}\Big\{ \gamma_\mu S_\beta(k)\gamma_\rho D_\beta^{\rho\sigma}(l) S_\beta(k+l) \qquad (7)$$
$$\times \gamma_\sigma S_\beta(k)\gamma_\nu S_\beta(p-k) \Big\},$$

where

$$\text{Tr}\Big\{ \gamma_\mu S_\beta(k)\gamma_\rho D_\beta^{\rho\sigma}(l) S_\beta(k+l) \gamma_\sigma S_\beta(k)\gamma_\nu S_\beta(p-k) \Big\}$$
$$= \chi_{\mu\nu}^{(1)} \quad ,$$
$$\chi_{\mu\nu}^{(1)} = \text{Tr}\Big[\gamma_\mu(\not{k}+m)\gamma_\rho(\not{k}+\not{l}+m)\gamma^\rho(\not{k}+m)\gamma_\nu(\not{p}+\not{k}+m) \Big]$$
$$= 8\Big[2g_{\mu\nu}m^2(m^2 - k.l)$$
$$+ 3m^2 \big\{ k_\mu(p-k)_\nu + k_\nu(p-k)_\mu \big\}$$
$$- g_{\mu\nu}k.(p-k)\big\{ 3m^2 - 2k^4 - 4k.l \big\}$$
$$- m^2 \big\{ l_\mu(p-k)_\nu + l_\nu(p-k)_\mu \big\}$$
$$+ g_{\mu\nu}l.(p-k)\big\{ m^2 - 2k^2 \big\} \Big] ,$$
$$\qquad (8)$$

while that vertex type correction to two loop photon self energy in figure 1 (b) can be written as

$$\Pi_{\mu\nu}^{(2)}(p) = e^4 \int \frac{d^4k}{(2\pi)^4} \int \frac{d^4l}{(2\pi)^4}$$
$$\text{Tr}\Big\{ \gamma_\mu S_\beta(k)\gamma_\rho D_\beta^{\rho\sigma}(l) S_\beta(k+l) \qquad (9)$$
$$\times \gamma_\nu S_\beta(k+l-p)\gamma_\sigma S_\beta(k-p) \Big\},$$

where

$$\text{Tr}\Big\{ \gamma_\mu S_\beta(k)\gamma_\rho D_\beta^{\rho\sigma}(l) S_\beta(k+l)$$
$$\gamma_\nu S_\beta(k+l-p)\gamma_\sigma S_\beta(k-p) \Big\} = \chi_{\mu\nu}^{(2)} \quad ,$$
$$\chi_{\mu\nu}^{(2)} = \text{Tr}\Big[\gamma_\mu(\not{k}+m)\gamma_\rho(\not{k}+\not{l}+m)\gamma_\nu$$
$$(\not{k}+\not{l}-\not{p}+m)\gamma^\rho(\not{k}-\not{p}+m) \Big]$$
$$= 8\Big[m^2 g_{\mu\nu}(2k^2 + l^2 + p^2 - m^2 \qquad (10)$$
$$- 2k.l - l.p - 2k.p) + 2m^2 l_\mu(p-k-l)_\nu$$
$$+ 2k.(k-p)\big\{ 2(k+l)_\mu(k+l)_\nu$$
$$- (k+l)_\mu p_\nu - (k+l)_\nu p_\mu - g_{\mu\nu}$$
$$[k^2 + 2k.l - k.p + l^2 - l.p] \big\} \Big] .$$

Due to manifest covariance in the theory, physically measurable couplings can be evaluated through contraction of vacuum polarization tensor $\Pi_{\mu\nu}$ with the metric in Minkowski space $g^{\mu\nu}$ [11]. The cold loops can then be integrated using the standard techniques of Feynman parametrization and dimensional regularization as discussed in the standard textbooks, whereas, in the same term, the evaluation of the hot loop after the cold one gives (see appendix)

$$g^{\mu\nu}\Pi_{\mu\nu}^{(1)+(2)}(p,T)=\frac{\alpha^2 T^2}{3}, \tag{11}$$

About reson why the rhs in eq. (11) is independent of p, we should explain that there are a number of terms depending on p but all are coefficients of (negligible) higher orders of T. Using eq. (6)

$$D=\frac{\alpha^2 T^2}{12}. \tag{12}$$

Thus, the photon "mass" is

$$m_{\text{"photon"}}=\sqrt{D}=\frac{\alpha T}{2\sqrt{3}}. \tag{13}$$

3. Radiative corrections on the speed of light

To keep gauge invariance of QED, instead of considering a "mass" for photon, we prefer to correct the speed of light; this is because it is in a "dispersive" medium (heat bath), which reduces its speed from 1 ($c=1$) to v. Thus, using eq. (13), the correction on photon energy is found as:

$$\Delta E=\frac{\frac{\alpha T}{2\sqrt{3}}}{\sqrt{1-v^2}}. \tag{14}$$

Indeed, we have considered this qualitative feature that the heat bath behaves like a dispersive medium for photon, because of screening effect due to vacuum polarization (virtual pairs of electron-positron), and thus reduces its speed. The energy correction due to scattering from the virtual electron-positron pairs should be of the order of m_e:

$$\Delta E \sim 2m_e. \tag{15}$$

Comparing eqs. (14) and (15):

$$\frac{\frac{\alpha T}{2\sqrt{3}}}{\sqrt{1-v^2}}\sim 2m_e, \tag{16}$$

or

$$v^2 \sim 1-\frac{\alpha^2 T^2}{48m_e^2}. \tag{17}$$

So, considering the approximation regime $T \ll m_e$, it is found that:

$$v \sim 1-\frac{\alpha^2 T^2}{96m_e^2}. \tag{18}$$

4. Discussion

In some papers, thermal effects on the speed of light have been already considered based on Euler-Heisenberg Lagrangian [8,10]:

$$\mathcal{L}_{eff}=-\frac{1}{4}F^{\mu\nu}F_{\mu\nu}-\frac{\alpha^2}{90\,m^4}\left[\left(F^{\mu\nu}F_{\mu\nu}\right)^2+\frac{7}{4}\left(F_{\mu\nu}\tilde{F}^{\mu\nu}\right)^2\right]; \tag{19}$$

The final results are a reduction in the speed of light proportional to the fourth order of temperature [12]

$$c=1-\frac{44\pi^2\alpha^2}{2025}\frac{T^4}{m^4} \tag{20}$$

One serious criticism of these calculations is that, applying the Euler-Lagrange equation to the starting Lagrangian, one cannot reach the standard wave equation but deals with a nonlinear equation by which we cannot introduce the "velocity" (speed) in its standard form:

$$-\frac{1}{c^2}\frac{\partial^2\vec{A}}{\partial t^2}\left[\frac{1+\frac{\alpha^2}{10m^4}\frac{1}{c^2}(\frac{\partial\vec{A}}{\partial t})^2}{1+\frac{\alpha^2}{30m^4}\frac{1}{c^2}(\frac{\partial\vec{A}}{\partial t})^2}\right]+\vec{\nabla}^2\vec{A}=0. \tag{21}$$

Some other papers [6, 9, 11, 13] have addressed the standard QED radiative corrections trying to find the electron charge renormalization, the electric permittivity, and the magnetic permeability in terms of temperature. Although there is no direct calculation of the speed of light in these papers, it can be simply shown that the resulting speed based on the calculated renormalized electron charge (and/or electromagnetic properties of the medium (heat bath)) is greater than 1 (c=1)! Considering no screening of magnetic fields due to the transverse nature of Π_{ij} implied by gauge invariance [14] and the following result from [9]:

$$Z_3=1+\frac{\alpha^2 T^2}{6m^2}, \tag{22}$$

it is found

$$v=\frac{1}{\sqrt{\varepsilon}}=\sqrt{Z_3}=1+\frac{\alpha^2 T^2}{12m^2} \tag{23}$$

Indeed, it seems the correct way of having a reasonable method and result is to consider the presence of the thermal heat bath as a dispersive medium and then have a kinematic calculation of the reduced speed of light due to its dynamically generated mass. Of course, as we have already mentioned, we don't consider the "mass" of photon as its rest mass (or any other similar quantity) to prevent the possibility of destroying gauge invariance of the theory.

Appendix

At low temperature, the hot contribution comes from the photon background only. Therefore, for example, eq. (7) for figure 1 (a) is simplified to

$$\Pi_{\mu\nu}^{(1)}=\Pi_{\mu\nu}^{(1)}(p,T=0)-\frac{2\pi ie^4}{(2\pi)^8}\int d^4k\int d^4l\frac{\chi_{\mu\nu}^{(1)}\delta(l^2)n_\beta(l)}{[k^2-m^2+i\varepsilon]^2[(k+l)^2-m^2+i\varepsilon][(p-k)^2-m^2+i\varepsilon]}, \tag{A-1}$$

then

$$g^{\mu\nu}\Pi^{(1)}_{\mu\nu}(p,T) = -2\pi i e^4 \int \frac{d^4k}{(2\pi)^4} \int \frac{d^4l}{(2\pi)^4} \times \frac{g^{\mu\nu}\chi^{(1)}_{\mu\nu}\,\delta(l^2)\,n_\beta(l)}{[k^2-m^2+i\varepsilon]^2[(k+l)^2-m^2+i\varepsilon][(p-k)^2-m^2+i\varepsilon]} \tag{A-2}$$

where

$$g^{\mu\nu}\chi^{(1)}_{\mu\nu} = 8[(2m^4 - m^2 l.p)$$
$$+ 3m^2(p_\mu + l_\mu)k^\mu + (-3m^2 + 2l.p)k^2 \tag{A-3}$$
$$+ 2k^2(k.p - k.l) - 2k^4]$$

Integrating over the cold loop using Feynman parametrization and dimensional regularization, one gets

$$g^{\mu\nu}\Pi^{(1)}_{\mu\nu}(p,T) =$$

$$\frac{e^4}{16\pi^5}\int_0^1 dx \int_0^x dy \int_0^y dz \int d^4l\;\delta(l^2)\,n_\beta(l)$$

$$\left[\frac{1}{[m^2 + 2l.pz(y-z)]^2}\right.$$

$$\left\{2m^4 - m^2(l.p)\Big[1 + 3y - 6z + 6z^2 - 6yz\Big]\right.$$

$$\left.-2(l.p)^2\Big[yz - z^2 + 4z^4 + 4z^2y^2 - 8z^3y\Big]\right\}$$

$$-\frac{1}{[m^2 + 2l.pz(y-z)]}$$

$$\left.\left\{-6m^2 + (l.p)\Big[1 - 12z^2 + 12yz\Big]\right\} - \frac{12}{\varepsilon}\right] \tag{A-4}$$

and

$$g^{\mu\nu}\Pi^{(1)}_{\mu\nu}(p,T) =$$

$$\frac{\alpha^2}{\pi^2}\int_0^1 dx \int_0^x dy \int_0^y dz \int_{-1}^1 d(\cos\theta)\int d|\vec{l}|\,\frac{|\vec{l}|}{e^{\beta|\vec{l}|}-1}$$

$$\times\left\{2m^4\left[\frac{1}{a_+^2} + \frac{1}{a_-^2}\right] - m^2\,|\vec{l}\,|[1 + 3y - 6z + 6z^2 - 6yz]\right.$$

$$\left[\frac{p_0 - |\vec{p}|\cos\theta}{a_+^2} - \frac{p_0 + |\vec{p}|\cos\theta}{a_-^2}\right]$$

$$-2|\vec{l}\,|^2\,[yz - z^2 + 4z^4 + 4z^2y^2 - 8z^3y]$$

$$\left[\frac{(p_0 - |\vec{p}|\cos\theta)^2}{a_+^2} + \frac{(p_0 + |\vec{p}|\cos\theta)^2}{a_-^2}\right]$$

$$+6m^2\left[\frac{1}{a_+} + \frac{1}{a_-}\right] - |\vec{l}\,|\Big[1 - 12z^2 + 12yz\Big]$$

$$\left.\left[\frac{p_0 - |\vec{p}|\cos\theta}{a_+} - \frac{p_0 + |\vec{p}|\cos\theta}{a_-}\right] - \frac{12}{\varepsilon}\right\} \tag{A-5}$$

where

$$a_\pm = m^2 \pm 2z(y-z)|\vec{l}\,|\,(p_0 \mp |\vec{p}|\cos\theta)$$

It should be noted that the 1/3 contribution is cancelled by the counter term diagram �—◯—. After a lengthy calculation of the integrals and regardless of the higher powers of T, one finds

$$g^{\mu\nu}\Pi^{(1)}_{\mu\nu}(p,T) = \frac{8\alpha^2 T^2}{9}. \tag{A.6}$$

Similarly

$$g^{\mu\nu}\Pi^{(2)}_{\mu\nu}(p,T) = -\frac{5\alpha^2 T^2}{9}, \tag{A.7}$$

then

$$g^{\mu\nu}\Pi^{(1)+(2)}_{\mu\nu}(p,T) = \frac{\alpha^2 T^2}{3}. \tag{A.8}$$

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